

Results from KamLAND

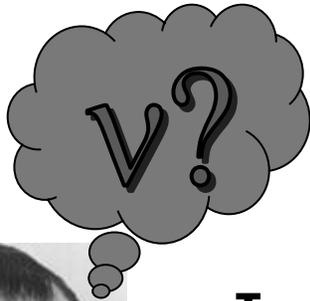
and future prospects

R.Svoboda, LSU

ANL Workshop on Trends in Neutrino Physics, 2003

Introduction

- A little neutrino history
- The Solar Neutrino Problem
- The KamLAND concept
- 1999-2002 Construction
- Backgrounds – almost none!
- Signals – reactor event sample
- Neutrino oscillations
- Conclusions and Deep Thoughts



The Heroic Age



In 1930 Pauli proposed the existence of a neutral fermion with very small mass to explain beta decay electron energy spectra

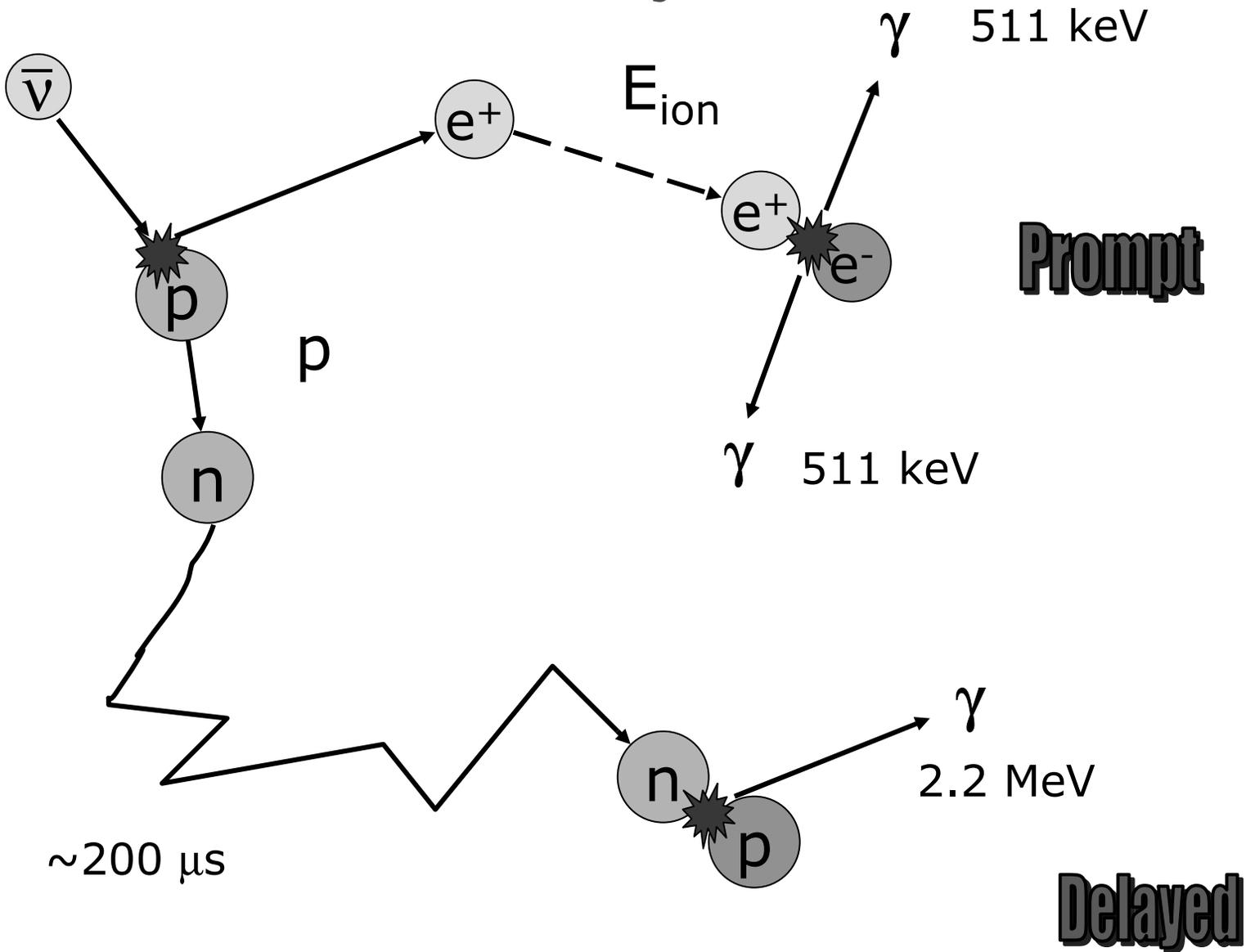
He feared such a particle might never be detected due to its elusive nature

1956 – Neutrinos Really Exist!



- Reines and Cowan detect neutrinos coming from the core of a nuclear reactor
- Photomultiplier tubes
- Liquid scintillator
- “high speed” coincidence circuits

Inverse Beta Decay



A Mirror of the Neutrino Spectrum

$$E_{\text{prompt}} = E_{\nu} - 0.78 \text{ MeV} - E_n$$

E_n : neutron recoil energy (small)

0.78 MeV: neutron – proton Δmc^2

Prompt energy spectrum is close to being original neutrino spectrum convoluted with the cross section!

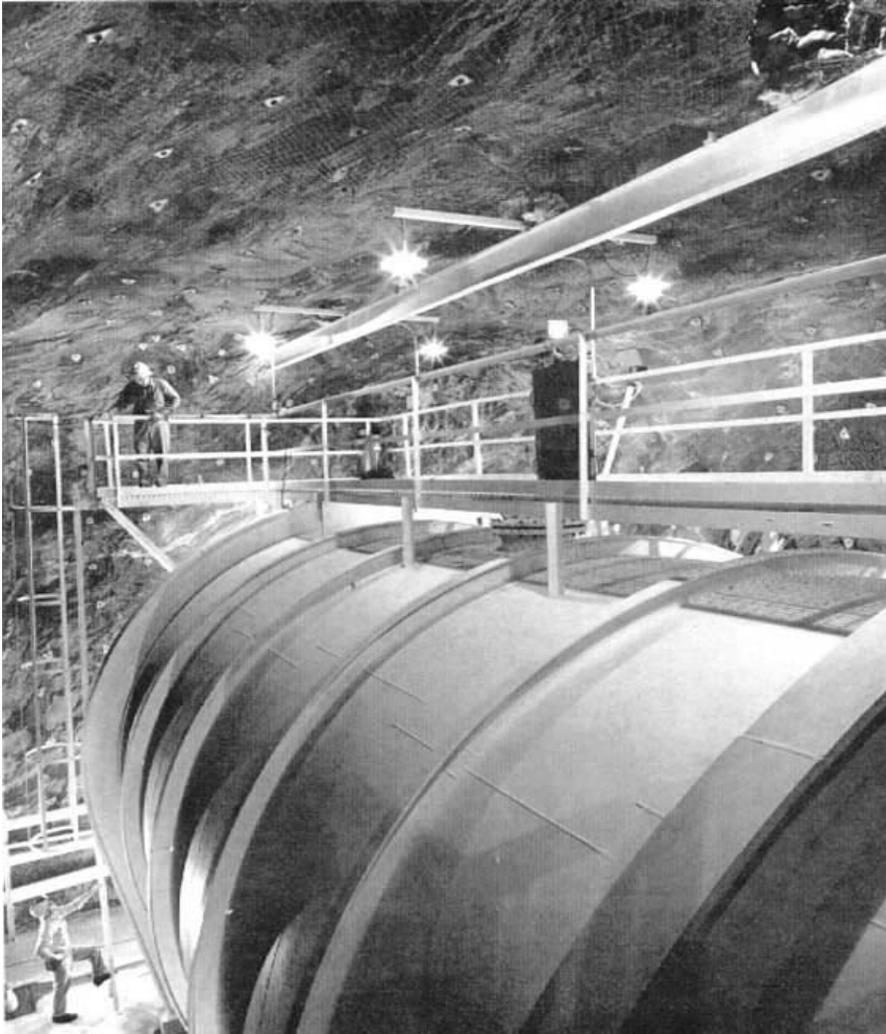
This double coincidence technique was the key to reducing the background to a low enough level to finally make the neutrino detectable

big people



small detector

Physics where the Sun Doesn't Shine...

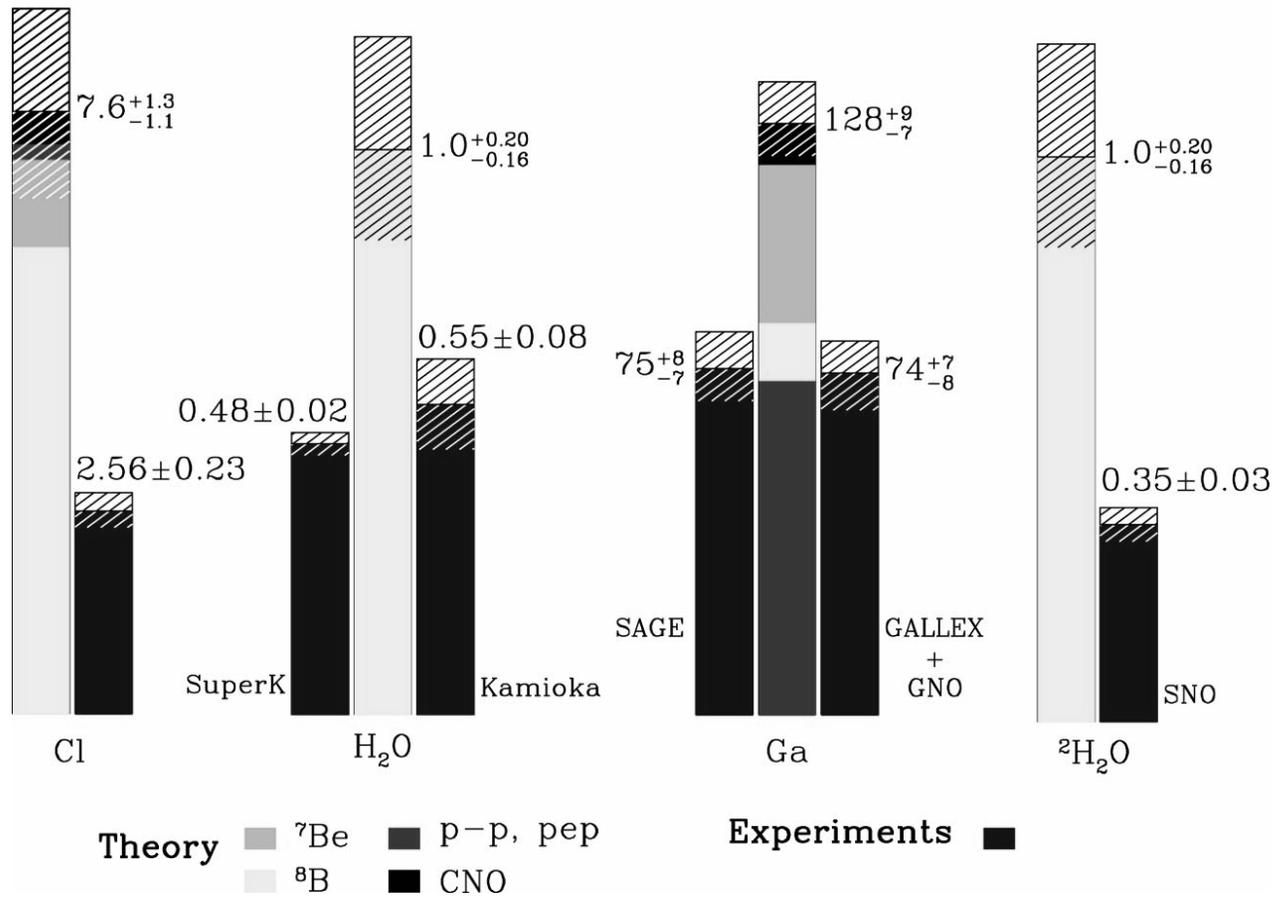


- after a failed attempt to detect the neutrino in 1954 Davis looks for solar neutrinos
- in 1968 he reported that there is a problem...
- there are less than 1/3 of the expected number of ν_e 's

Fast Forward 40 years...

Too Few Solar ν 's

Total Rates: Standard Model vs. Experiment
Bahcall-Pinsonneault 2000



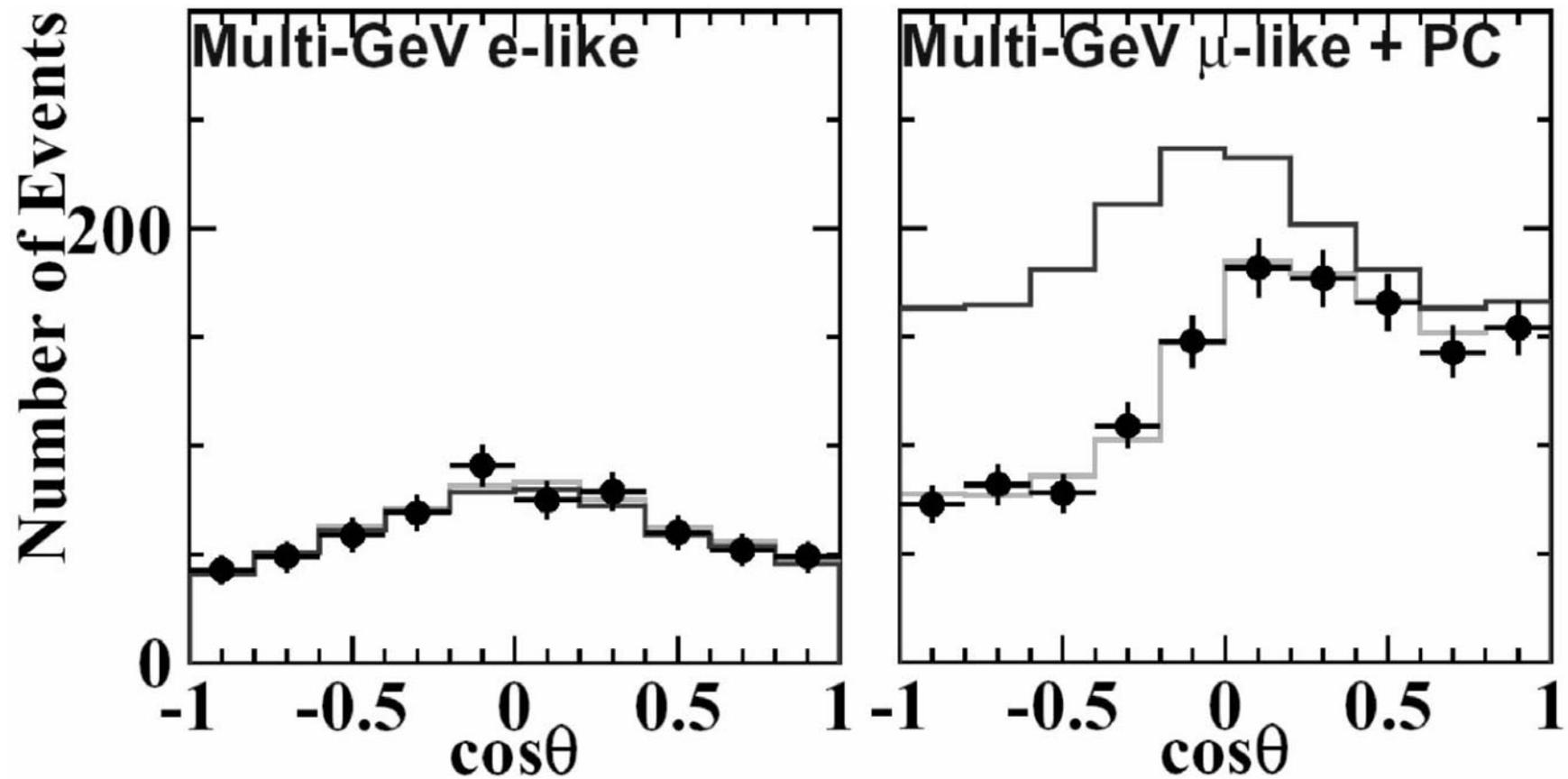
Neutrino Oscillations

$$|v_{\mu,t}\rangle = |1\rangle \cos \theta e^{-im_1^2 t / 4p} + |2\rangle \sin \theta e^{-im_2^2 t / 4p}$$

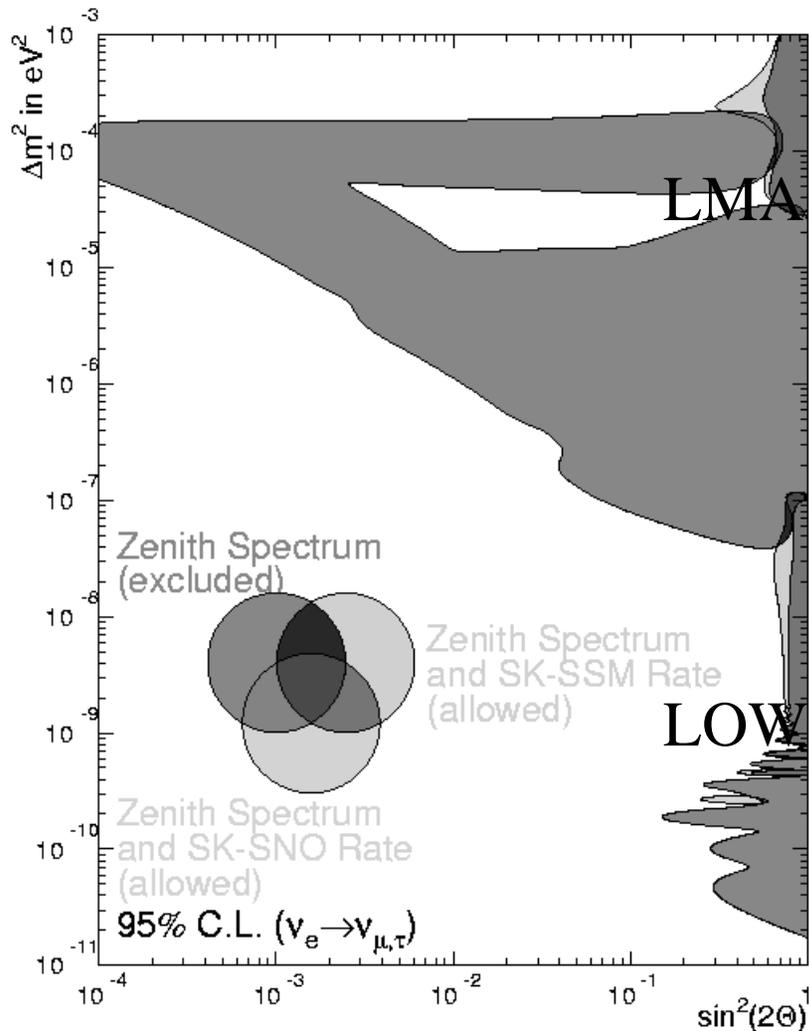
- “survival probability” for two components
- $|\langle v_{\mu} | v_{\mu} \rangle|^2 = 1 - \sin^2 2\theta \sin^2[1.27 \delta m^2 (L/E)]$

a very sensitive way to look
for neutrino mass

1998: Super-Kamiokande Neutrino Oscillation Measurements Survive Longer than the “New Economy”



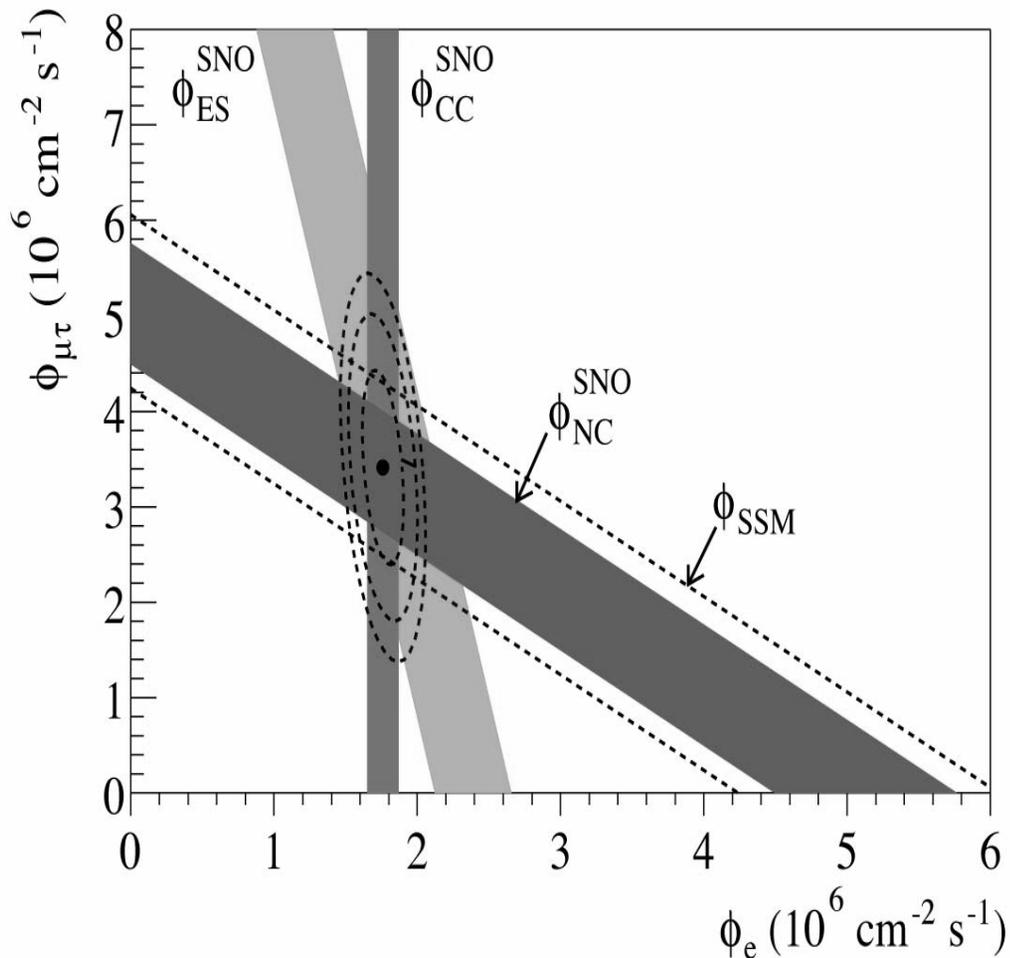
Super-K Solar ν 's



No spectral distortion or “matter effects” seen in propagation through the earth

- Two major solutions left:
 - LMA
 - LOW

SNO measures total neutrino flux via NC capabilities



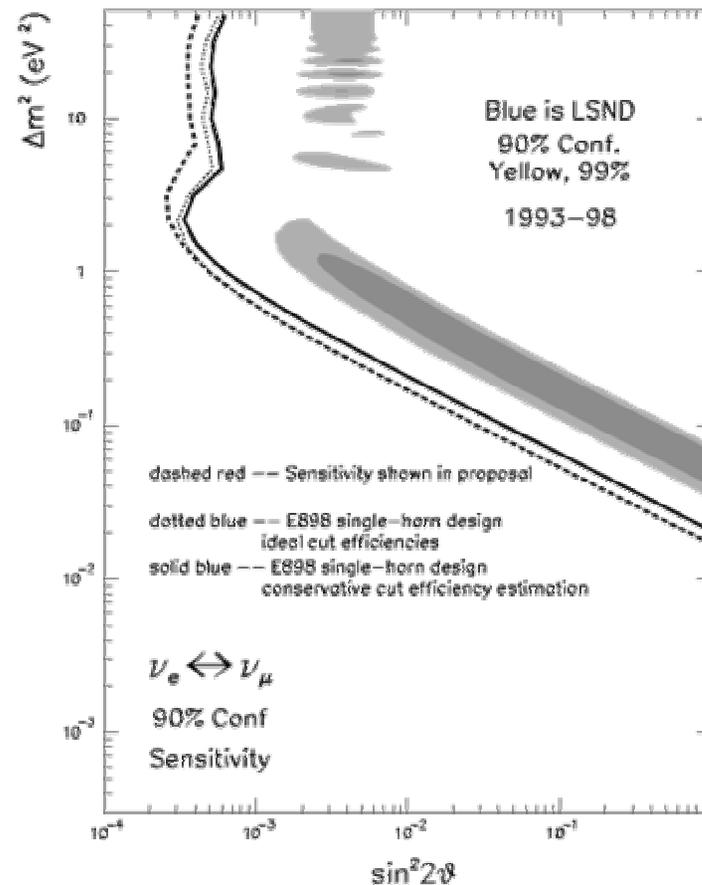
total flux matches expectations!

combining all experiments now leaves only one oscillation region left

LMA

Unanswered Questions

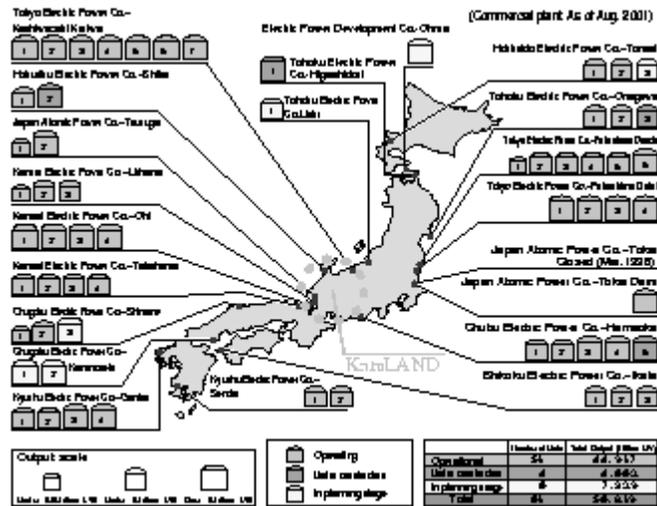
- Is LMA the solution to the Solar Neutrino Problem? (excluding other solutions not a very satisfying way to say LMA is correct)
- If so, what are the exact mixing parameters?
- Could $\sin^2 2\theta$ be very close to, or exactly 1?
- LSND? CPT violation? Do anti-neutrinos have the same mass hierarchy as neutrinos?



The KamLAND Concept

- re-do the Reines-Cowan experiment but with a BIG detector and MANY reactors at distances ~ 100 km or more
- this would allow testing the LMA solution to the Solar Neutrino Problem with a well-understood neutrino source
- JAPAN is a good place to do this

Japan is an excellent location



1 km deep in
Kamioka Mine
1



water-filled
active veto

buffer oil
region

1325 17"
554 20" PMT's

scintillator
filled balloon
6.5 m radius
1000 tons

液体シンチレータ反ニュートリノ
観測施設設置区域
(改造前)

カミオカンデ

平成9年度

電源供給室



実験室区域入口

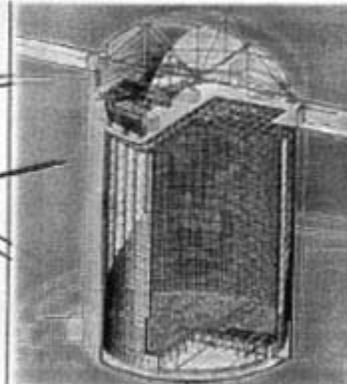


坑道分岐点

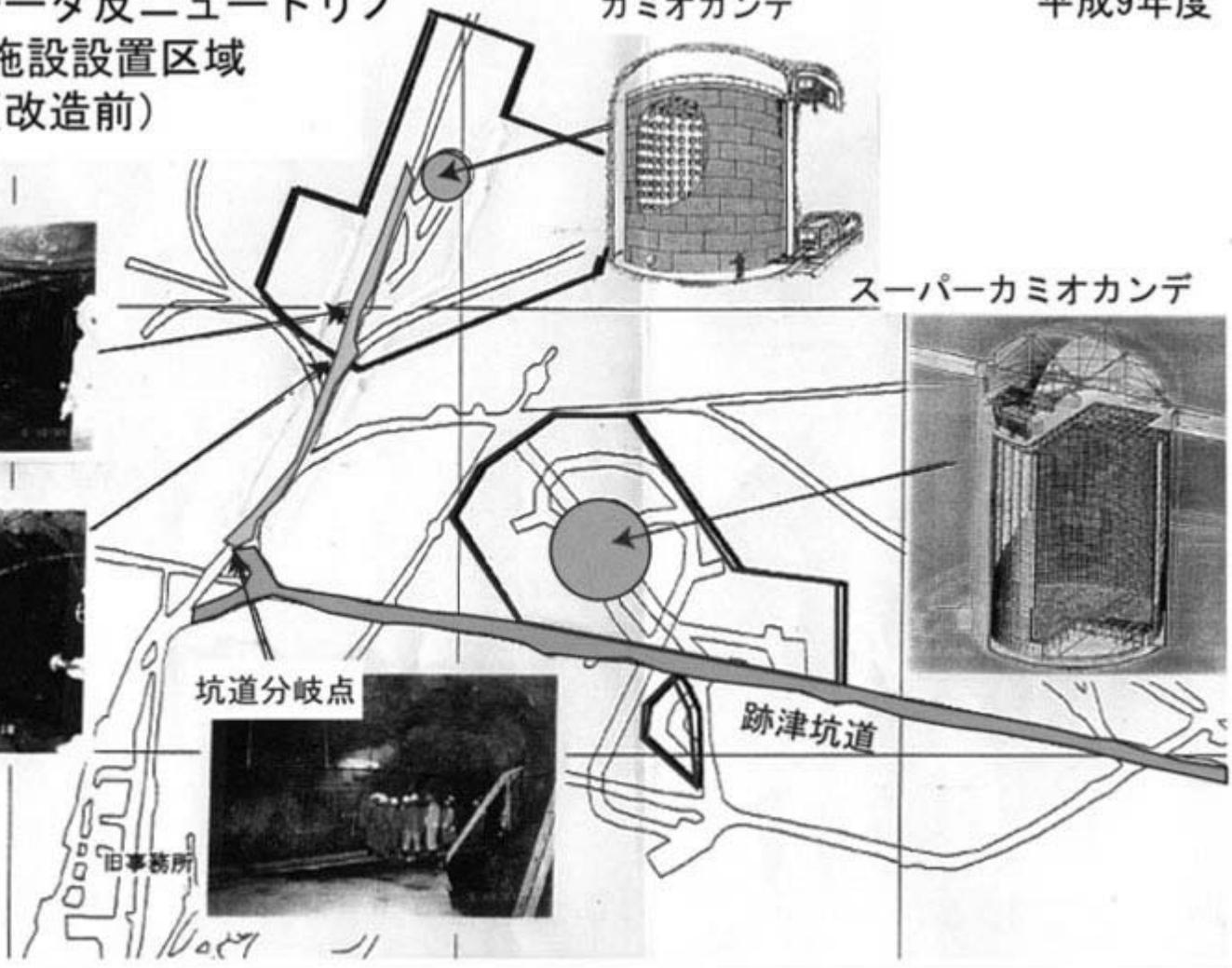


旧事務所

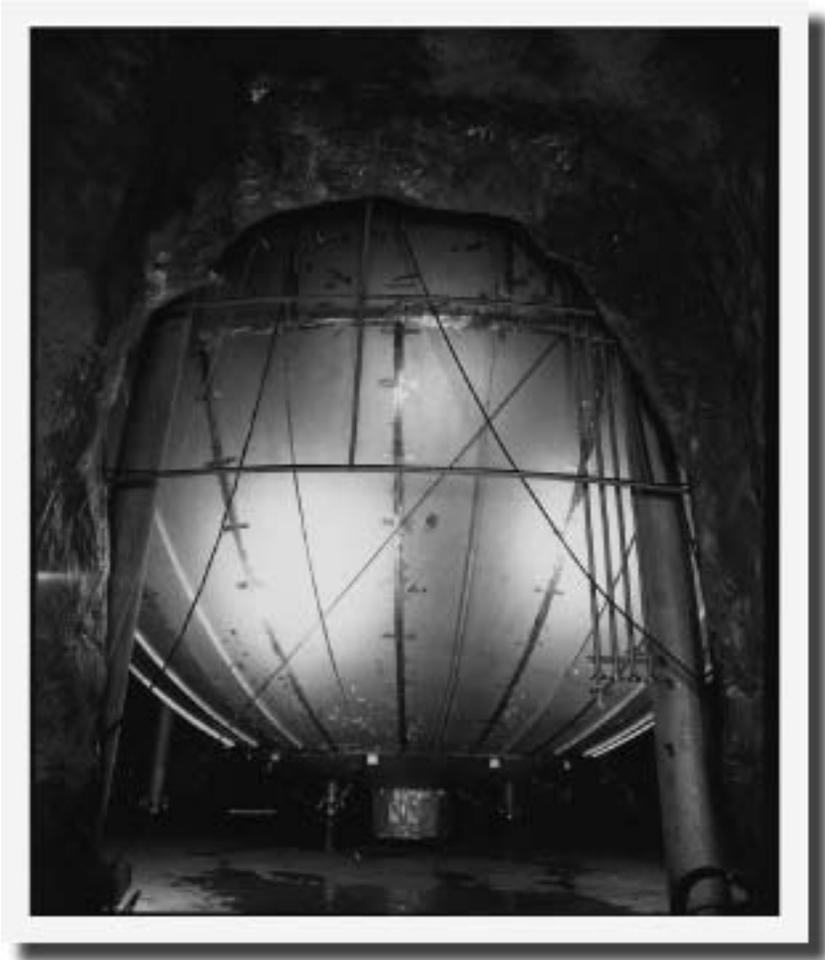
スーパーカミオカンデ



跡津坑道



KamLAND Construction

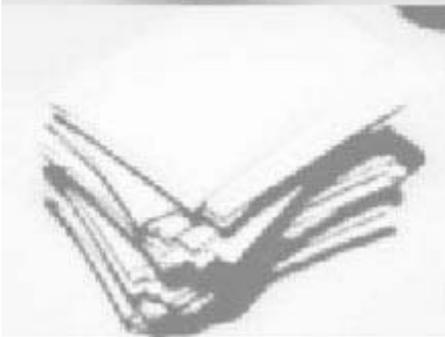


- use Kamiokande site
- cavity renovation started in 1998
- PMT installation 2000
- Oil filling 2001
- taking data January, 2002
- whew!

Expected Signal

- Reactor anti-neutrino flux as a function of time and distance
- energy threshold and calibration
- fiducial volume resolution
- efficiency of event selection cuts
- live time





Livetime

- March 4, 2002 to October 6, 2002
- 145.1 live days
- uncertainty 0.07%

Trigger Efficiency

- 200 hits ~ 0.7 MeV (prompt)
- 120 hits ~ 0.4 MeV (delayed)
- electronics deadtime $\ll 0.1\%$
- For reactor neutrinos, we use a data reduction threshold of 0.9 MeV and an analysis threshold of 2.6 MeV
- for this analysis the efficiency is essentially 100% for both prompt and delayed triggers

Energy Calibration

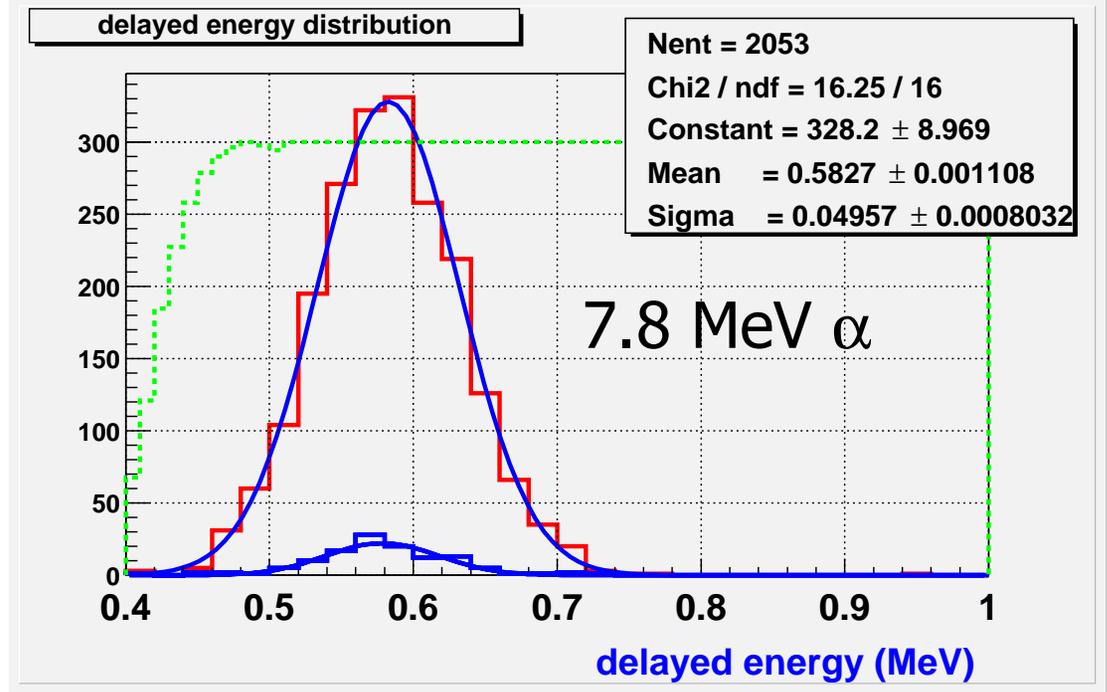
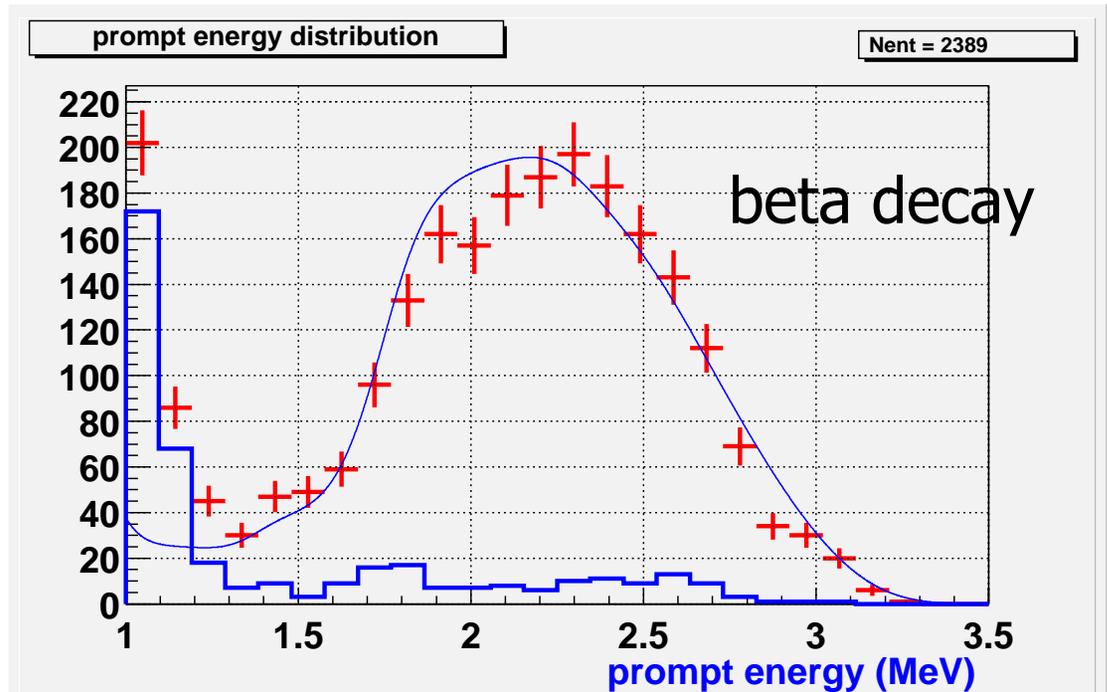
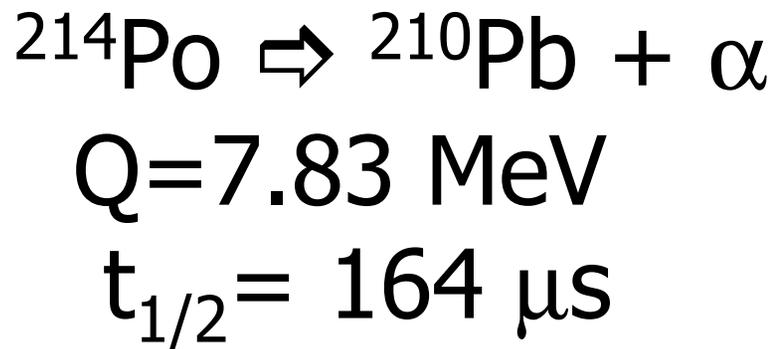
- ^{60}Co , ^{65}Zn , ^{68}Ge , and AmBe sources provide γ lines in the range 0.5 to 7.6 MeV
- these are deployed along the z-axis on a regular basis
- about 300 p.e./MeV are observed, giving an energy resolution of 7.5% at 1 MeV.
- off the z-axis the energy reconstruction is confirmed using muon spallation products and contaminants in the detector

Energy Reconstruction

- “Standard” corrections for absorption, PMT acceptance, PMT wall density, scattering from detector walls, ropes, etc are all made
- In addition to sources we have UV lasers and LED’s to do these
- ...but KamLAND is more complicated than the “typical” water Cherenkov detector

Living in the Material World

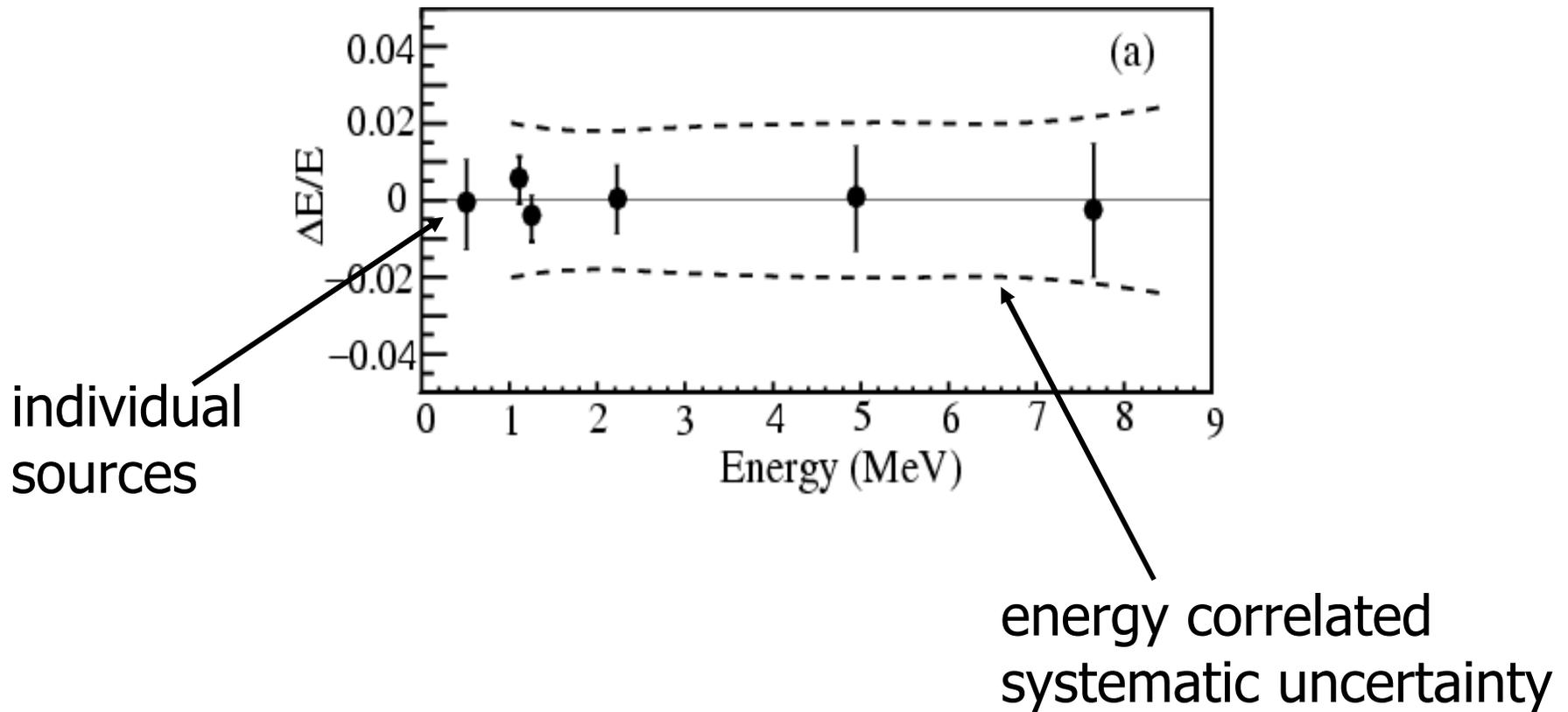
- liquid scintillator has a non-linear response in energy due to saturation effects. This is on the order of a few percent for γ 's to factors of 10-15 for highly-ionizing α 's
- Cherenkov light production is a significant contribution to the visible energy due to the PPO waveshifter in the scintillator
- We measure these effects *in situ* using calibration sources and “tagged” α 's from U/Th/Rn contaminants ^{212}Po and ^{214}Po



Energy Reconstruction Uncertainty

- also use ^{40}K , ^{208}Tl , ^{12}N , ^{12}B , and thermal neutron capture from muon spallation to test off-axis reconstruction
- variation within the fiducial volume less than 0.5%
- energy scale time variation is less than 0.6%
- total energy scale systematic uncertainty at 2.6 MeV is 1.91%
- uncertainty in rate above 2.6 MeV due to energy calibration uncertainty:
2.13%

Results of Energy Calibration



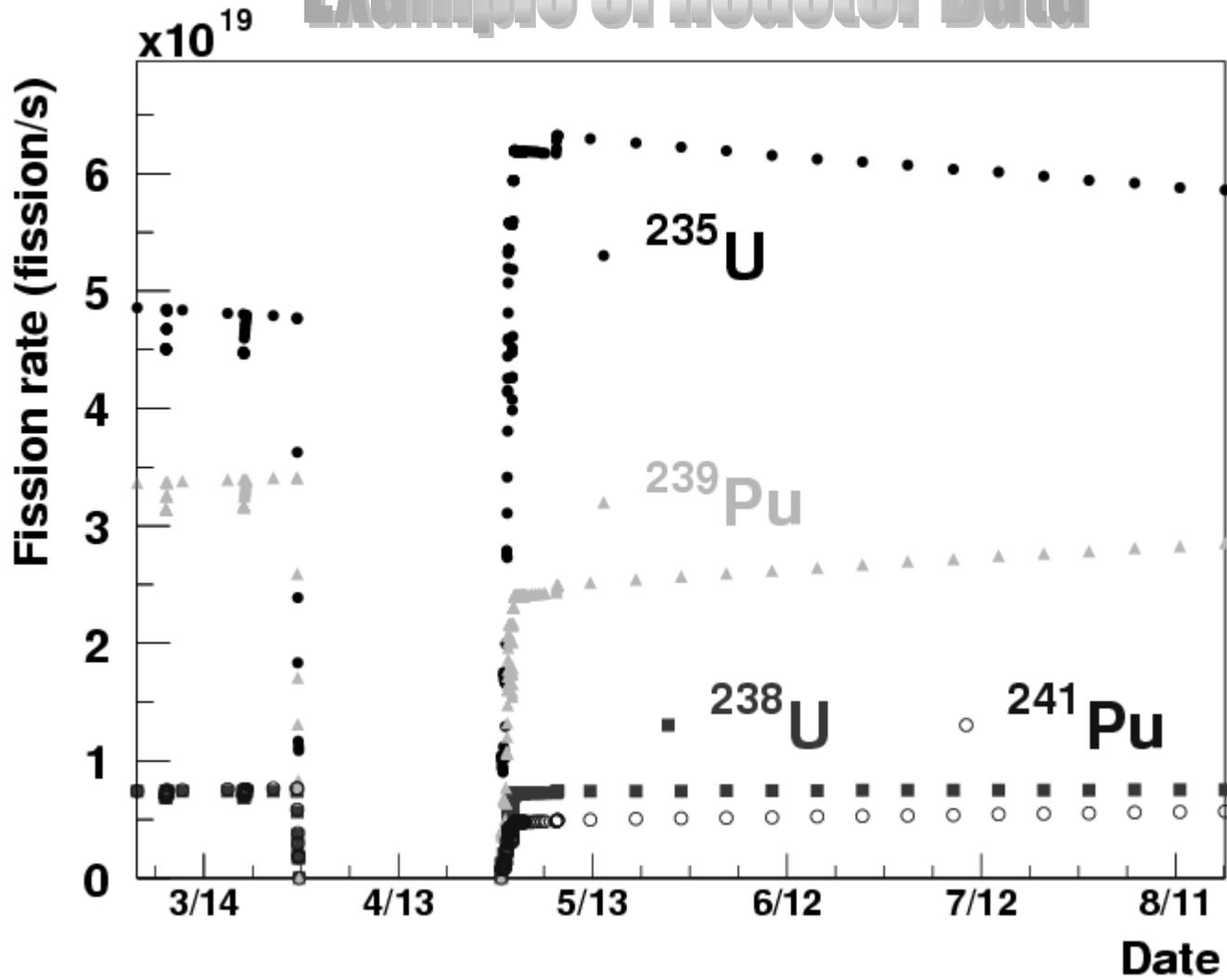
Reactor Thermal Power

- Thermal power history provided for all Japanese reactors. By law, this is required to be known to better than 2%.
- While all reactors considered, 79% of the flux comes from 26 reactors in the distance range 138-214 km
- Another 6.7% from one reactor at 88 km
- All other reactors more than 295 km away
- Korean reactors 2.5% (estimated to 10%)
- Rest of the world is 0.7% (estimated to 50%)

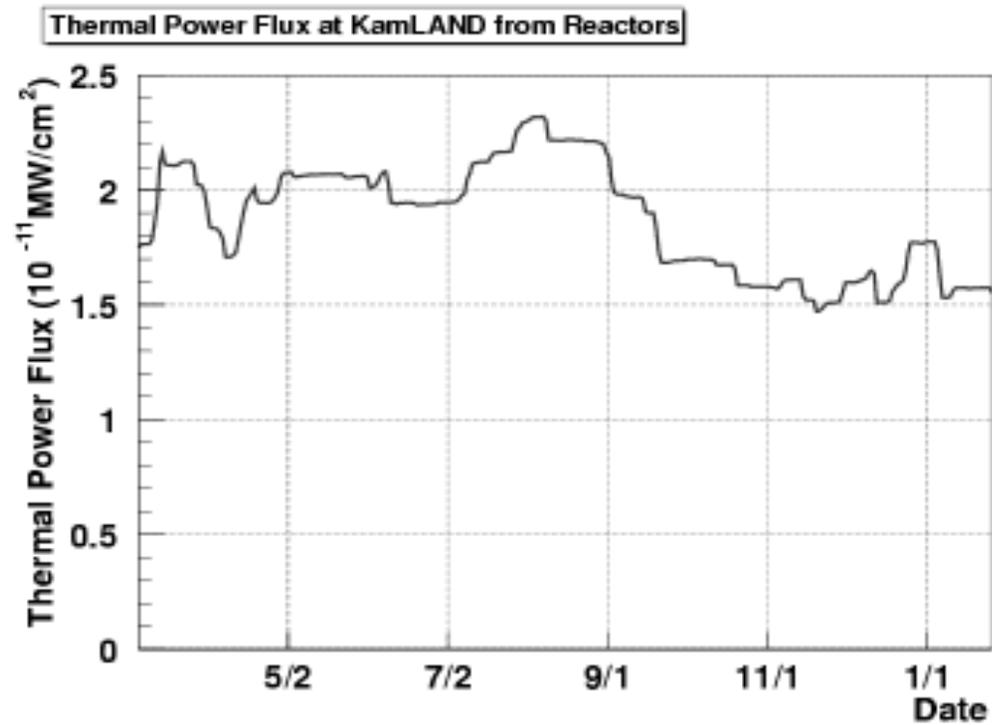
Reactor Fuel Loading and History

- ✓ Average fuel loading: ^{235}U : ^{238}U : ^{239}Pu : ^{241}Pu
is 0.568: 0.078: 0.297: 0.057
- ✓ error in anti-neutrino flux from uncertainty in fuel loading is less than 1%
- ✓ uncertainty in flux per fission is 2.5% based on estimates from several calculations
- ✓ power history uncertainty folded into the contribution of delayed beta emitters adds another 0.3% uncertainty

Example of Reactor Data



Time Variation of Summed Flux

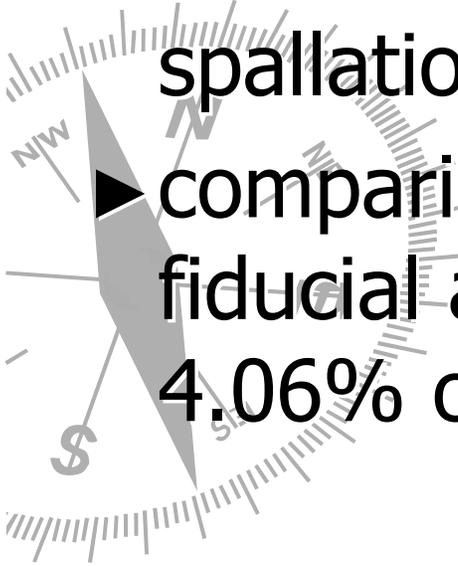


Uncertainty in the Interaction Rate

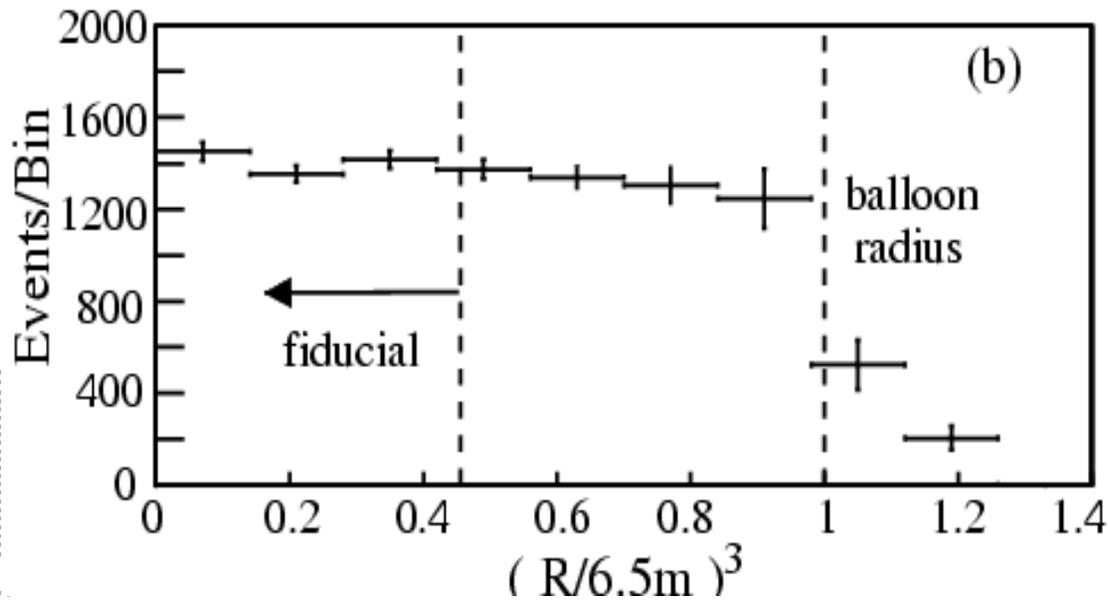
Reactor Power:	2.05%
Fuel Loading:	1.00%
History:	0.28%
Spectrum:	2.48%
Cross Section:	<u>0.20%</u>
	3.39%

Fiducial Volume Resolution

- ▶ The total mass of scintillator added to the detector is uncertain to 2.13%
- ▶ the ratio of fiducial to total volume is checked using neutron capture from muon spallation.
- ▶ comparison of the capture rate inside the fiducial and total volume agrees to within 4.06% of the expected volume ratio.



Vertex Distribution of neutron capture from muon spallation events agrees to within 4.06% of expectation



total uncertainty in target mass: 4.58%

Event Selection

- Energy and fiducial volume cut
- muon spallation cut (loss of 11.4% in live time)
- time between prompt and delayed event ($0.5 < \Delta T < 660 \mu\text{s}$)
- distance between prompt and delayed vertex cut ($\Delta L < 1.6\text{m}$)
- distance from central axis thermometer array $> 1.2 \text{ m}$
- total selection efficiency: $78.3(1.6)\%$

- the total efficiency of these cuts was verified using the absolute source strengths of the ^{60}Co and ^{65}Zn sources
- the total efficiency of the distance and time cuts was verified with AmBe source to within 1% using the 4.4 MeV gamma as a tag for the neutron de-excitation of $^{13}\text{C}^*$

Expected Event Rate

- Total systematic uncertainty in the expected rate is 6.42%
- expected rate: 86.8 ± 5.6 events above 2.6 MeV
- Background?

Accidental Coincidences

- Singles rate of 30 Hz with time and distance cuts
- expected number of accidentals is very small
- 0.0086 events

^8Li and ^8He

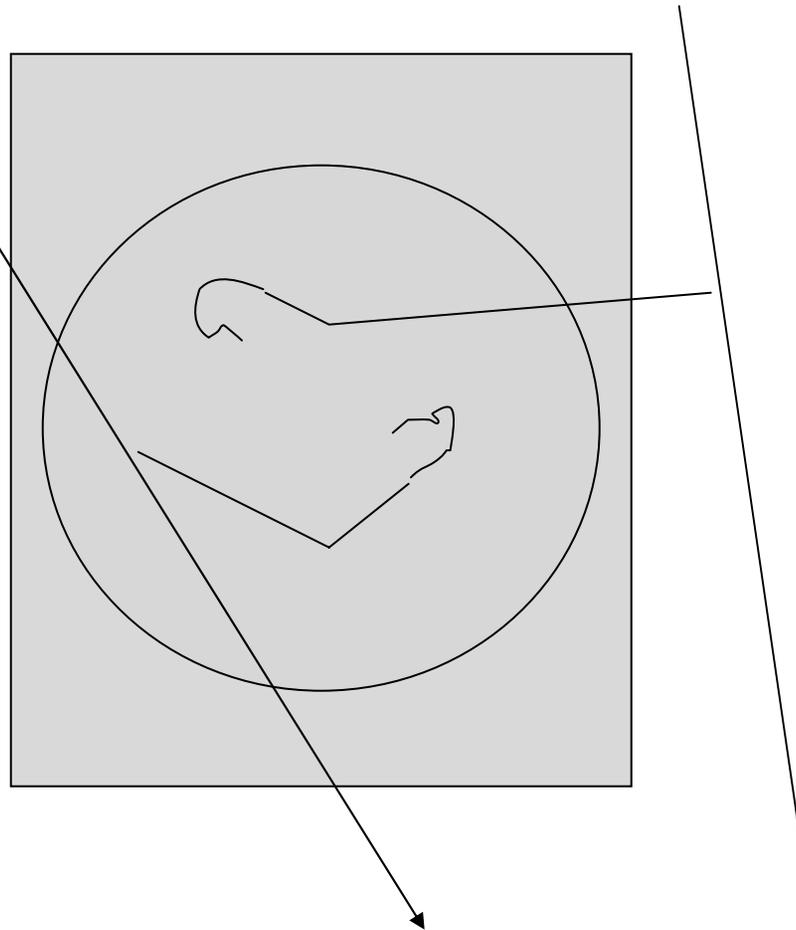
- These spallation products (produced by the passage of muons through the detector) can mimic reactor neutrino events in that they can neutron de-excite to a beta unstable daughter
- such backgrounds can be almost eliminated by applying a time and distance cut around muon induced events
- estimated background in sample 0.94 ± 0.85

Fast Neutrons

muons may produce
fast neutrons which
can mimic reactor
events

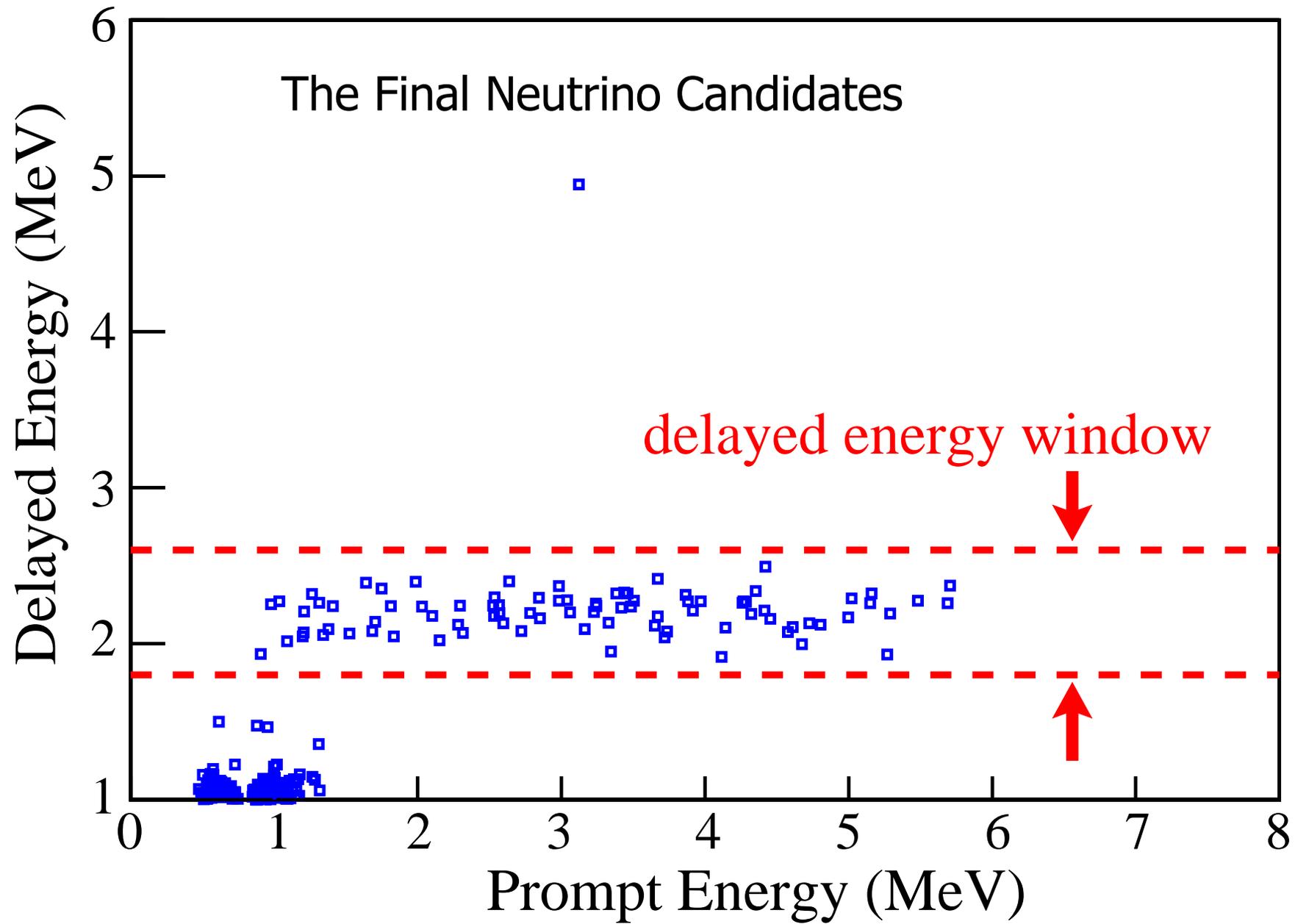
~3,000 neutrons/day

important calibration
source

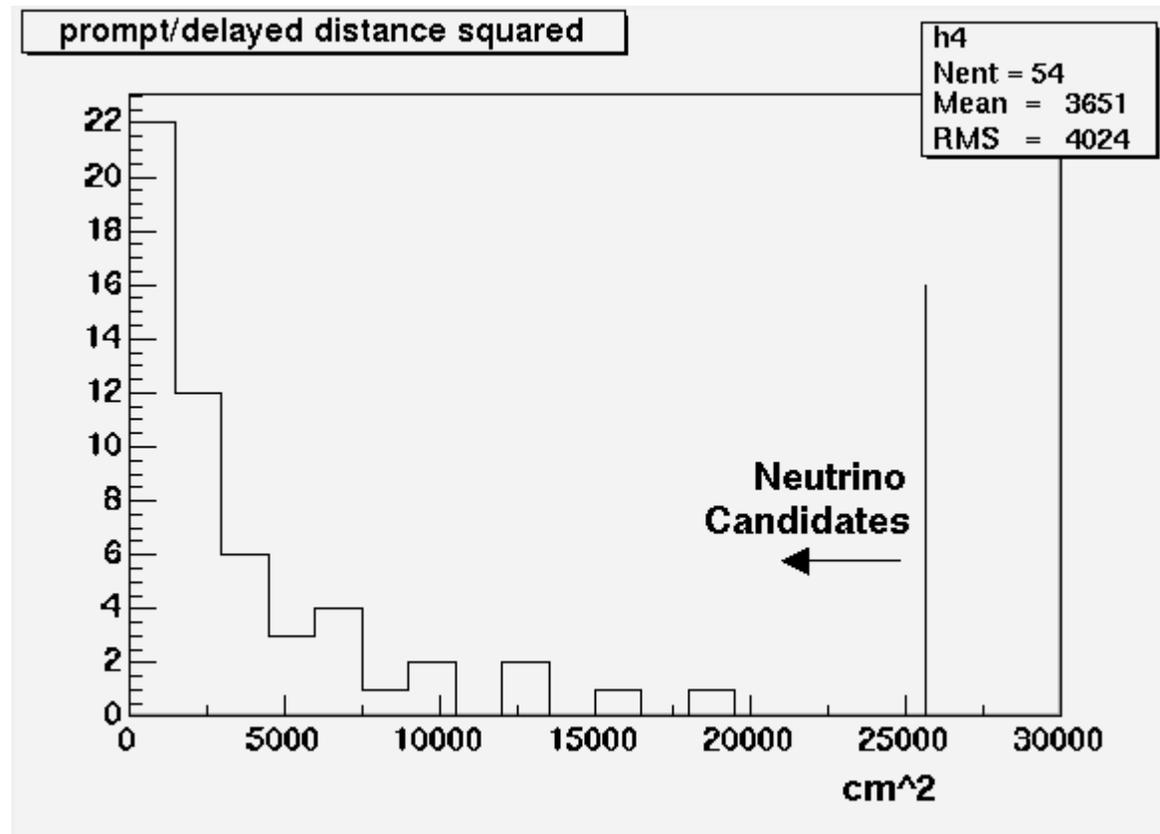


Fast Neutrons

- almost all neutrons are tagged by precursor muons in the detector or water veto
- we can measure the distance from the track of the capture to let us calculate the rate of missed captures from muons in the rock
- OD-only muons with a fast neutron let us measure the energy distribution of the recoil protons
- we estimate background above 2.6 MeV
<0.5 events



ΔL Distribution of neutrino events



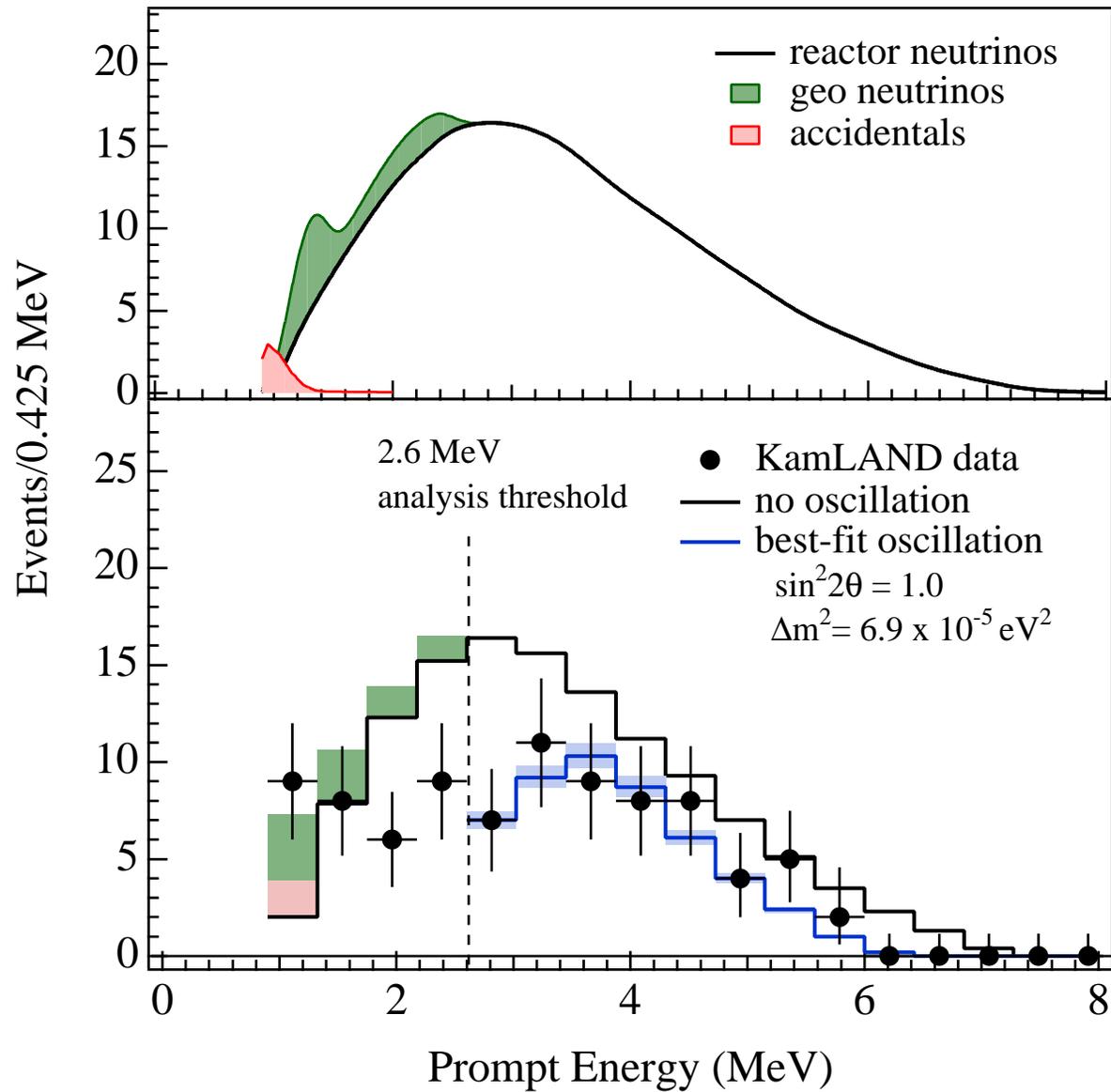
not sensitive to cut value

Measured Event Rate

- To avoid ambiguities associated with geo-neutrinos we make a cut at 2.6 MeV
- 54 events survive these cuts
- measured/expected ratio:

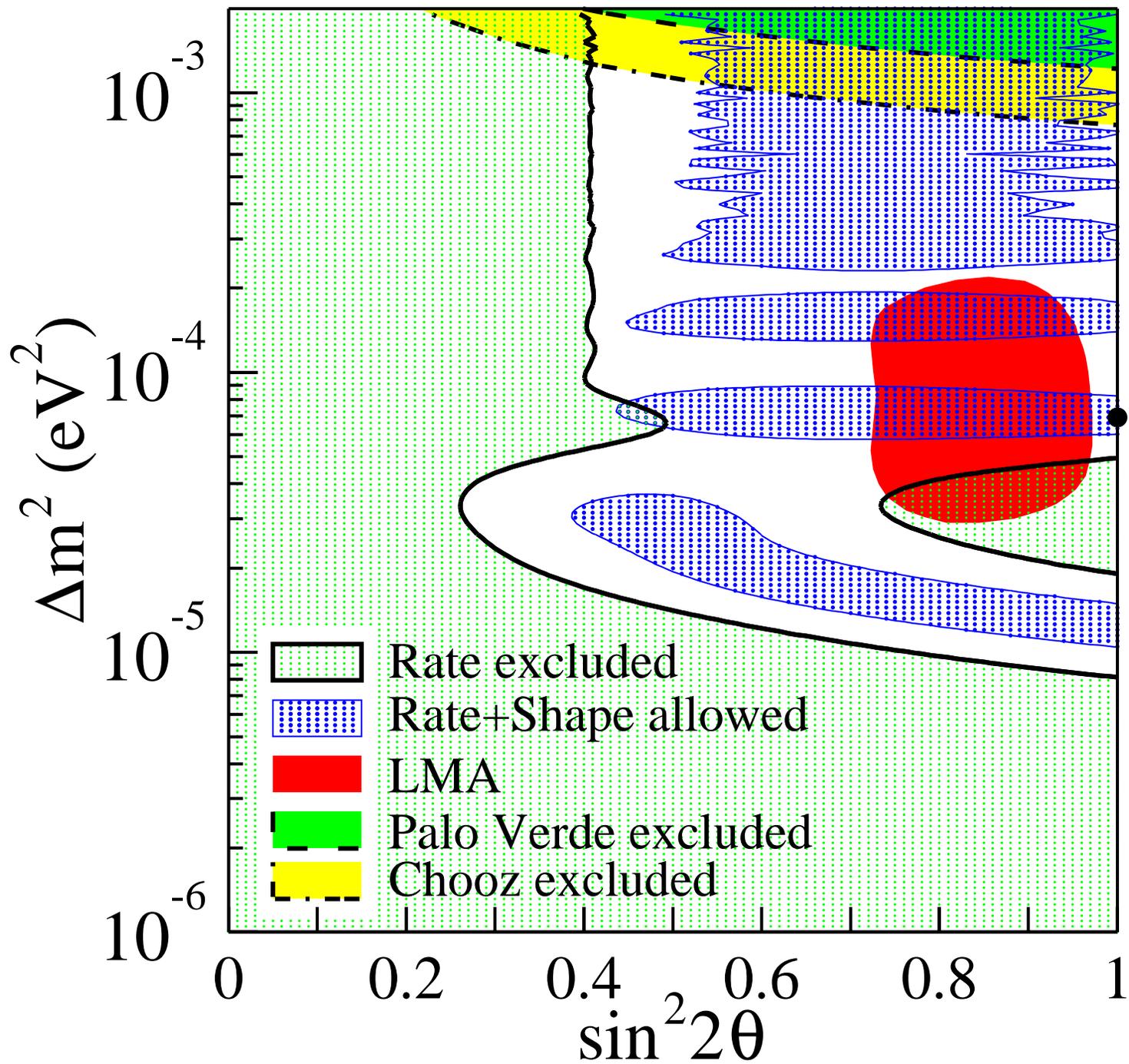
$0.611 \pm 0.085(\text{stat.}) \pm 0.041(\text{syst.})$
probability of no disappearance $< 0.05\%$

Measured Spectrum



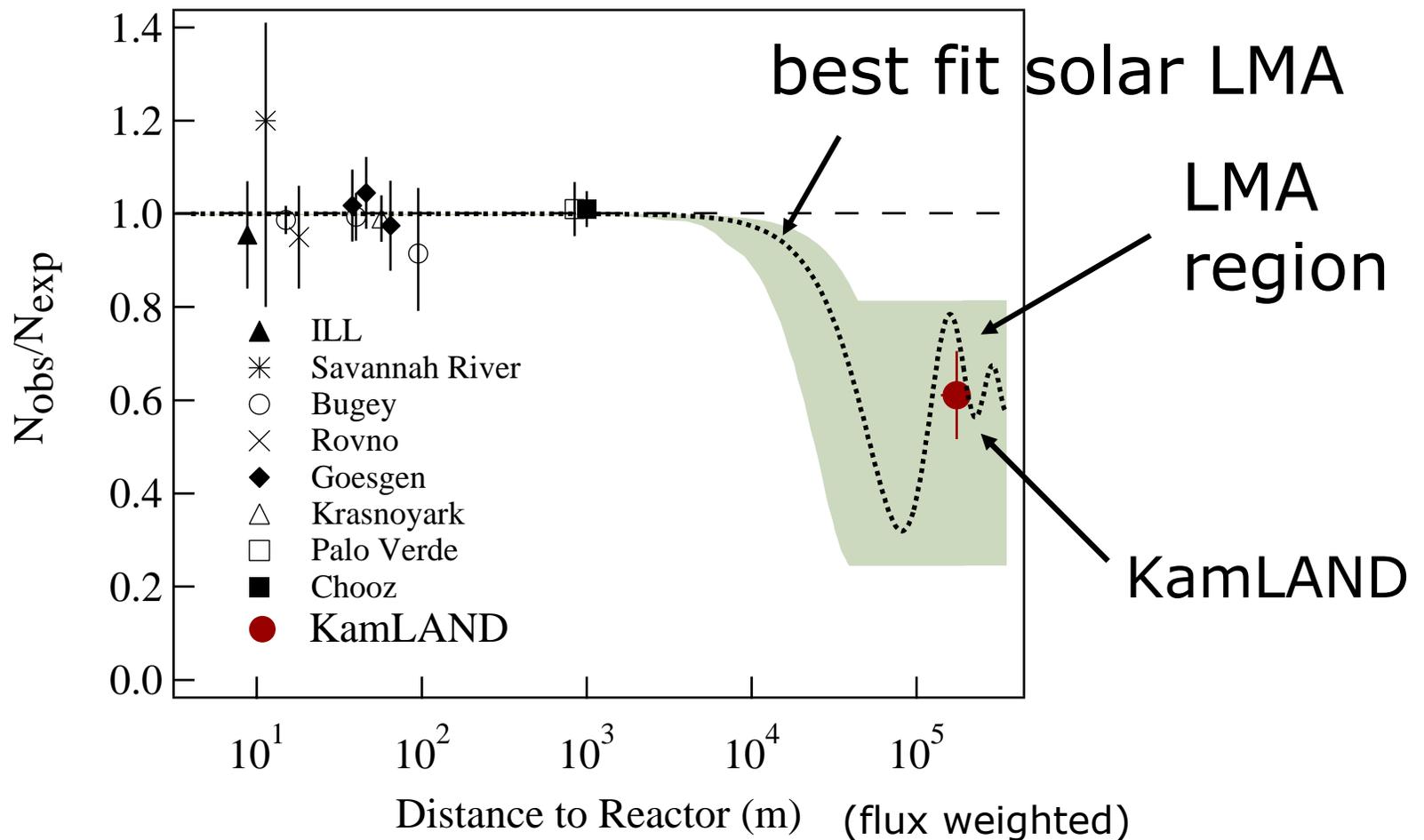
Neutrino Oscillations

- Fit using a two-component mixing model
- Simple χ^2 fit to the rate as a hypothesis test
- Maximum likelihood test using spectral shape and taking into account correlated uncertainties



- best fit: $\sin^2 2\theta = 1.0$ $\Delta m^2 = 6.9 \times 10^{-5} \text{ eV}^2$
- global minimum at $\sin^2 2\theta = 1.01$
- distorted spectrum consistent at the 93% confidence
- but renormalized undistorted spectrum also consistent at 53%

Comparison with Other Experiments



Conclusions

- A deficit of reactor neutrinos is seen at a 4.1σ level
- backgrounds are very low
- Combining spectral information with the simple rate measurement gives a two-component best fit of $\sin^2 2\theta = 1.0$ and $\Delta m^2 = 6.9 \times 10^{-5} \text{ eV}^2$
- best fit may (will?) likely change with additional spectral data

Deep Thoughts: *The Future*

- 3 live years: statistical error 13.9% to 5.0%
- improve energy calibration
- improve fiducial volume resolution (new calibration arm)
- improve energy resolution (increase light collection by 40%)
- systematic error 6.4% to $\sim 4.7\%$ or less
- error in rate goes from 15.3% to 6.9%

$\sin^2 2\theta$

- suppression factor $f \sim 1 - \sin^2 2\theta (1/2)$ gives $\sin^2 2\theta = 0.78(0.19)$
- a 24% measurement using rate only
- note that spectrum mostly improves δm^2 resolution
- estimated improvements give $\sim 10\%$ measurement
- *Veritas filia temporis* (Francis Bacon)
- ^8B solar neutrinos, geo-neutrinos, nucleon decay results on the way
- ^7Be neutrinos will take some time
- thanks!