A deep space photograph showing a vast field of stars. A prominent bright star in the lower-left quadrant exhibits a significant diffraction pattern, with multiple rays of light extending outwards. The rest of the sky is filled with numerous smaller, distant stars of varying colors and brightness.

What Cosmology/Astrophysics and Neutrino Physics Can Teach Each Other

Working Group Leaders:
Steve Barwick and John Beacom

Participants

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Vince Cianciolo

Paul Langacker

Jon Thaler

Mike Dragowsky

John LoSecco

Neil Weiner

Perspective

"If [there are no new forces] ---- one can conclude that there is no practically possible way of observing the neutrino." Bethe and Peierls, Nature (1934)

Cosmological

- Big-bang nucleosynthesis consistency
- Neutrino hot dark matter models ruled out

Astrophysical

- Neutrinos from SN 1987A observed
- The solution of the solar neutrino problem

Major Topics

- Leptogenesis
- BBN
- Dark energy
- Dark matter
- WIMP detection
- UHE neutrinos
- SN neutrinos



- Number of flavors
- Sterile neutrinos
- Dirac vs. Majorana
- Mass scale
- Mixing parameters
- Cross sections
- Exotic properties

Neutrinos are a key to
New physics in the Universe
New physics beyond the Standard Model

Working Group Goals

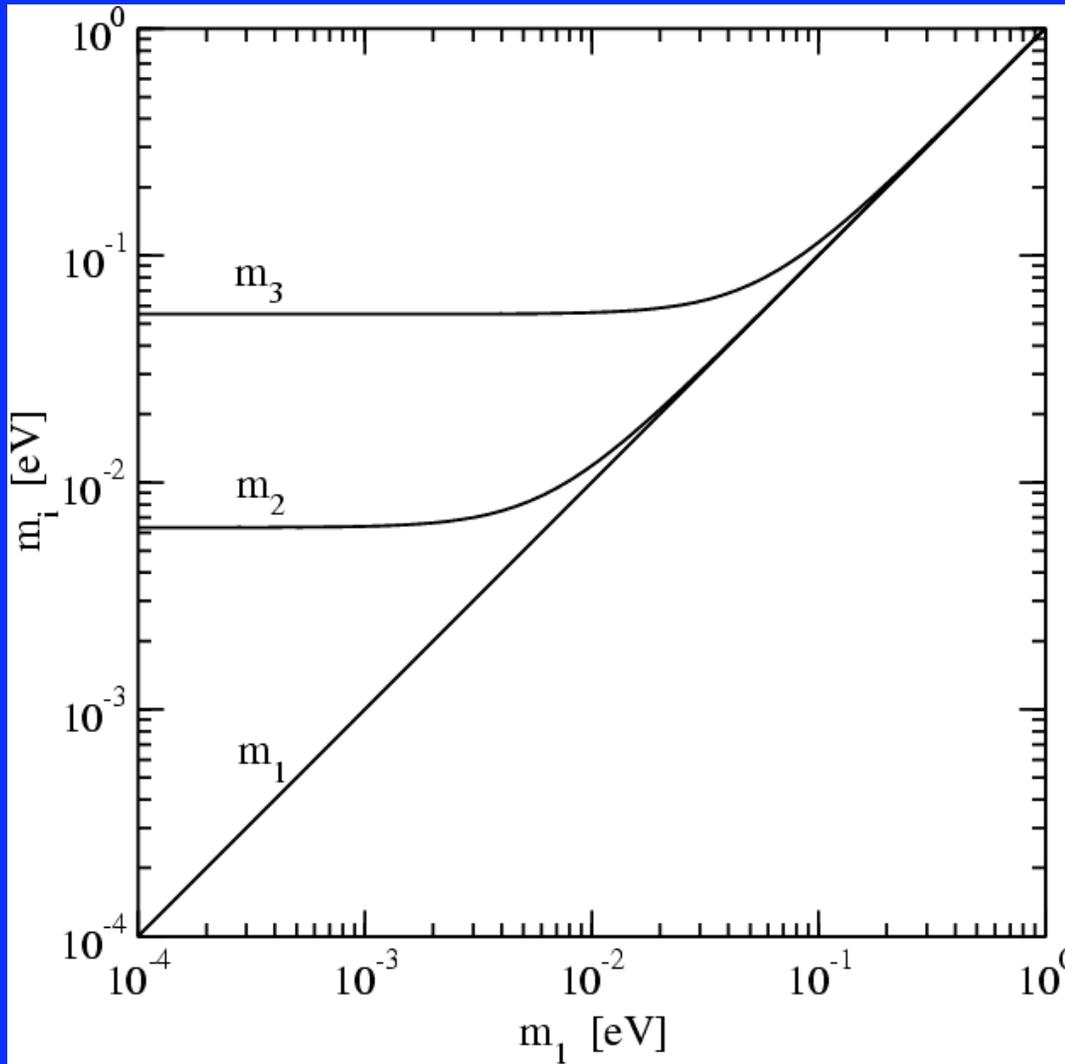
Four guiding goals:

1. New experiments in neutrino astrophysics
2. Added value to cosmological observations
3. Key role of theory in making connections
4. Strong connections to other working groups and nuclear/particle laboratory data

Selected Key Opportunities

- Discovery of neutrino mass using cosmological data
 - Key hint for model-building
 - Guide and foil for beta / double beta experiments
- Discovery of astrophysical neutrinos
 - Unique probe of extreme environments
 - Unique probe of neutrino properties, energy frontier
- Connections to new astrophysical/cosmological data
 - Detailed astrophysical models
 - Quest for identifying the particle dark matter
 - Fundamental theory towards the GUT scale

Neutrino Masses



Normal Hierarchy

$$m_1 = m_1$$

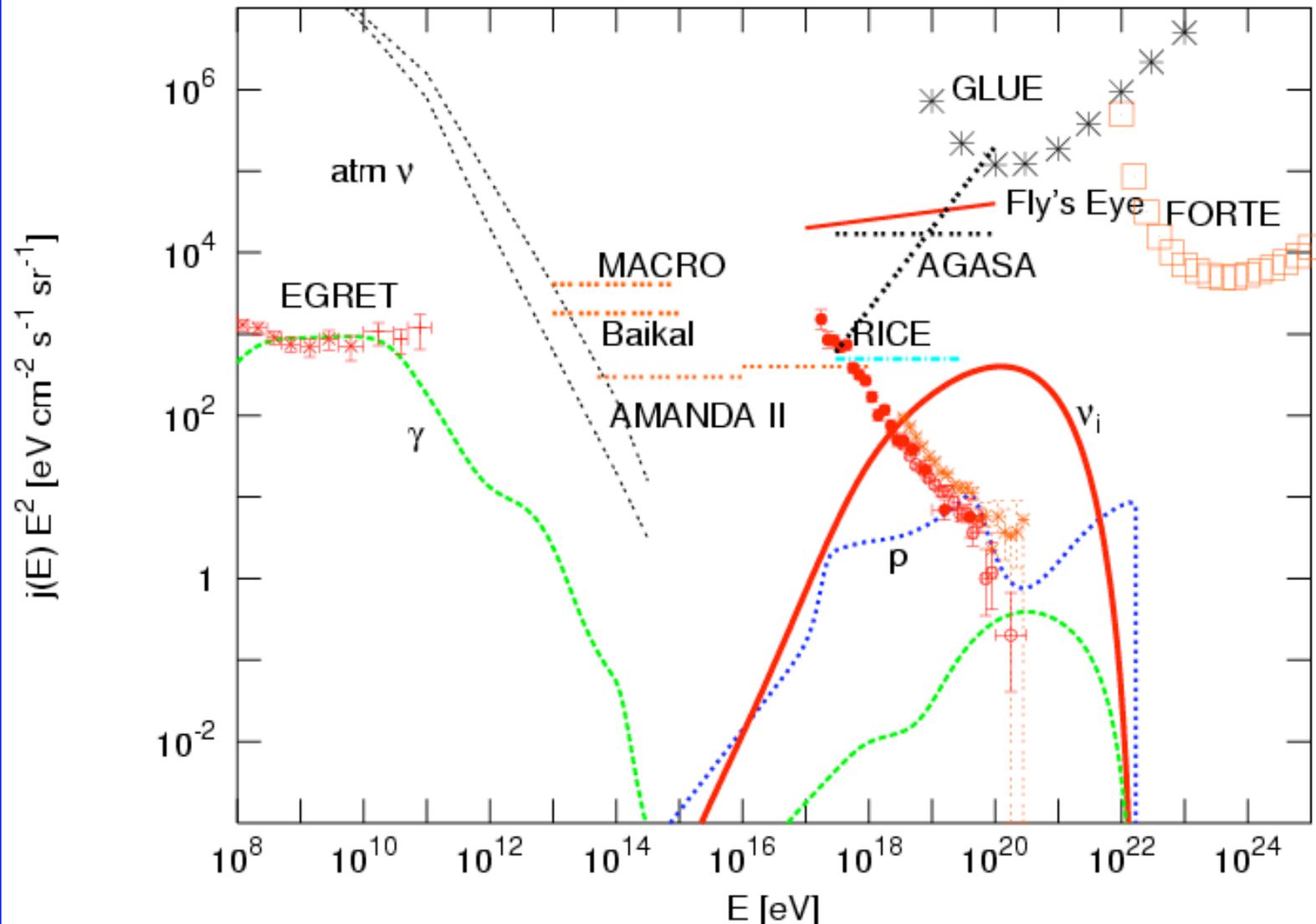
$$m_2 = \sqrt{m_1^2 + \delta m_{\text{solar}}^2}$$

$$m_3 = \sqrt{m_1^2 + \delta m_{\text{solar}}^2 + \delta m_{\text{atm}}^2}$$

$$\frac{m_3}{m_2} \leq \frac{\sqrt{\delta m_{\text{atm}}^2}}{\sqrt{\delta m_{\text{solar}}^2}} \leq 10$$

Beacom and Bell, PRD 65, 113009 (2002)

Existing Neutrino Limits



Semikoz, Sigl, hep-ph/0309328

Working Group Plans

- Have defined key topics and physics questions
- Have begun assembling working group
- Will provide updates through a web page
- Your participation and comments are welcomed
- Contact information:
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